

Ground based and satellite observations of the Ionospheric Alfven Resonance

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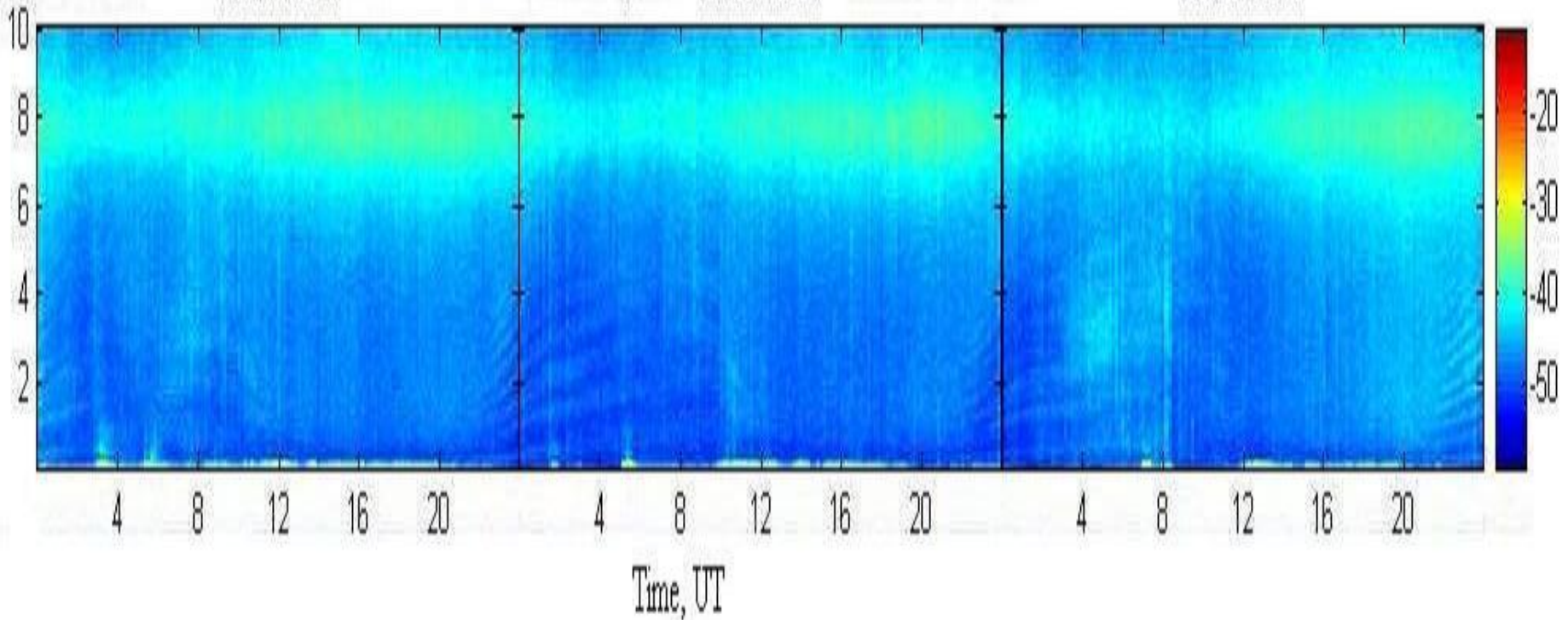
Institute of Radio Astronomy, NASU, Kharkov, Ukraine

Sample of the Spectral Resonance Structure of the Ionospheric Alfvén Resonance

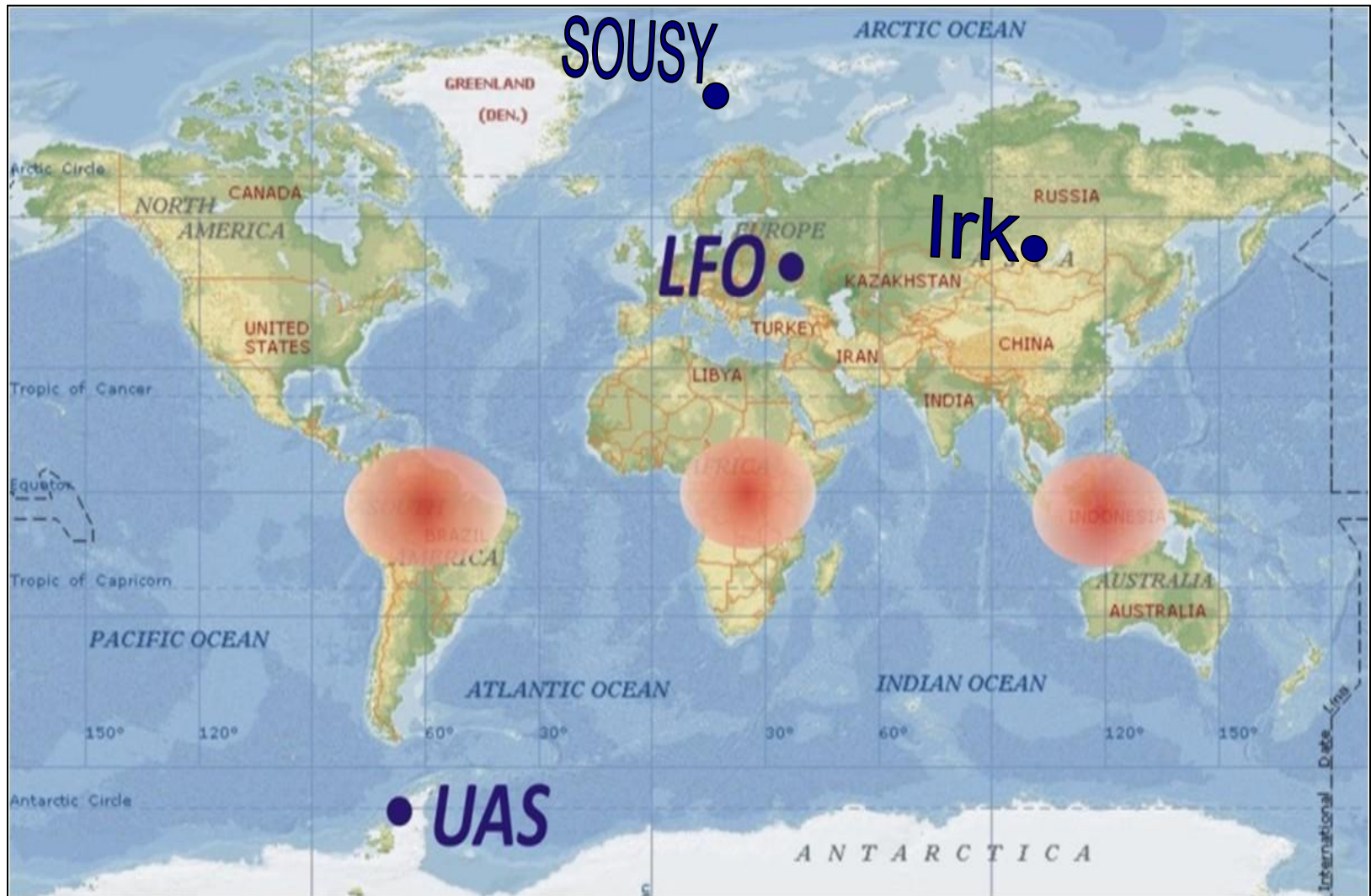
13.09.2006

14.09.2006

15.09.2006



Stations of observation



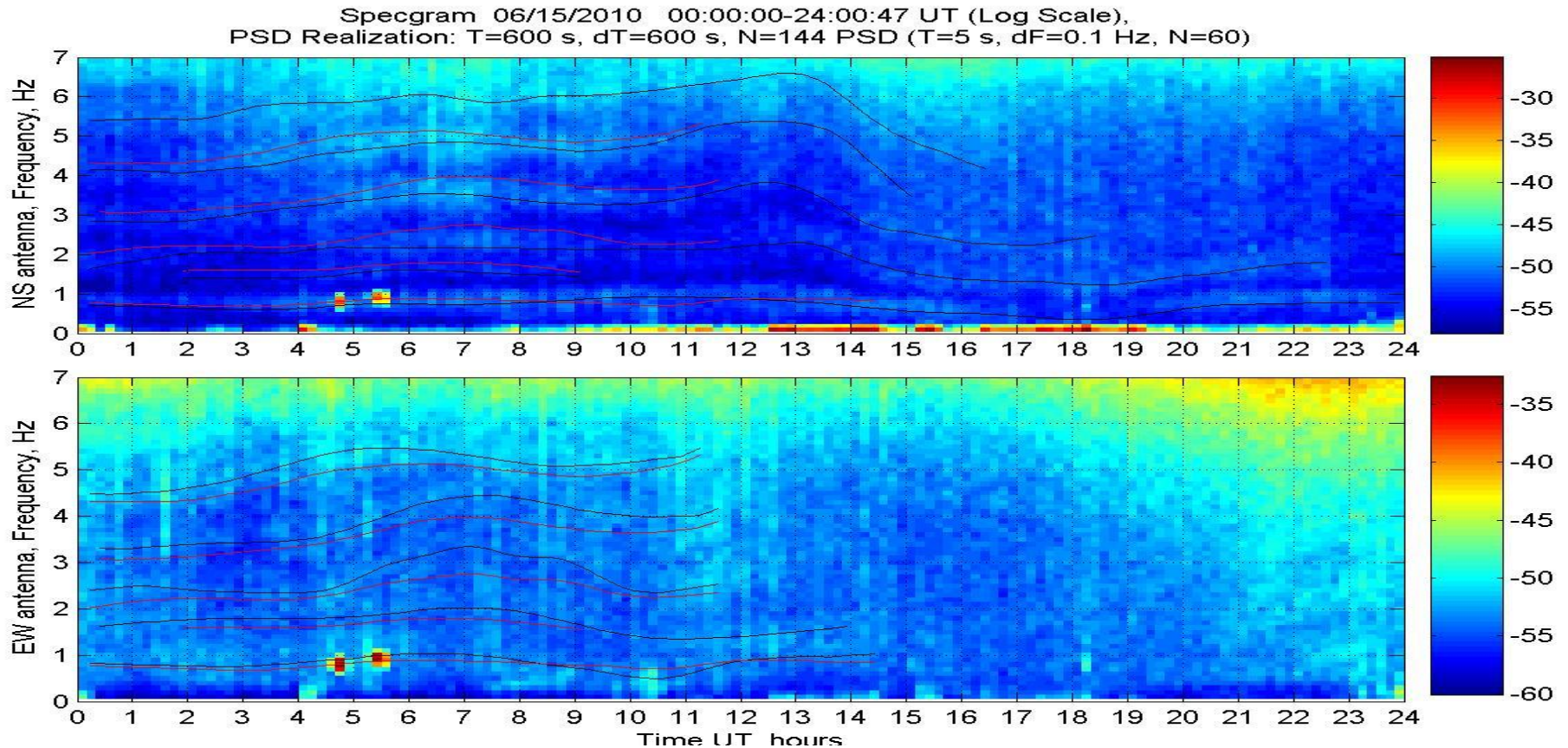
Equipment

- UAS
- SOUSY
Magnetometer Lemi112
Frequency range 0.001 – 80 Hz
- Sayan Solar Observatory
Magnetometer Lemi30
Frequency range 0.001 - 40 Hz
- Lviv Centre of Institute for Space Research NASU NSAU
<http://isr.lviv.ua>

- LFO IRA NASU
(<http://ri.kharkov.ua/geospace/en/observatory/elf.html>)

Frequency range	0.5 - 40 Hz
Sampling rate	125 Hz
Retransmitter frequency	433 MHz
Retransmitter power	1 mW
Retransmitter range	≈ 3 km
Supply voltage (Autonomous part)	6 V

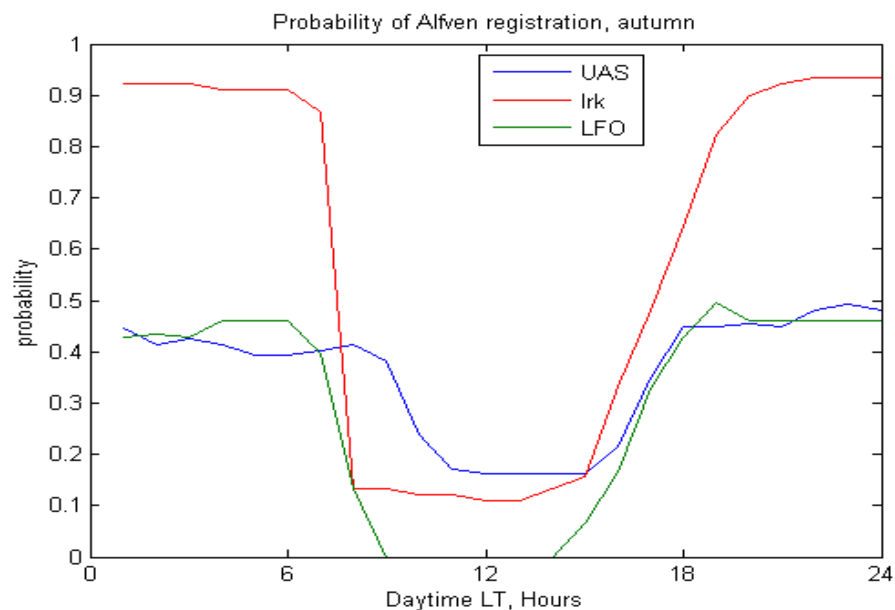
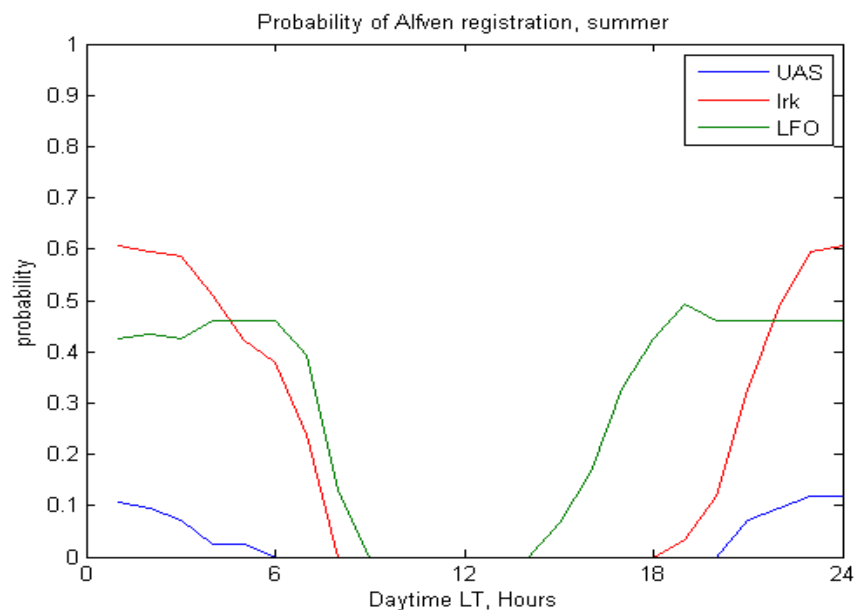
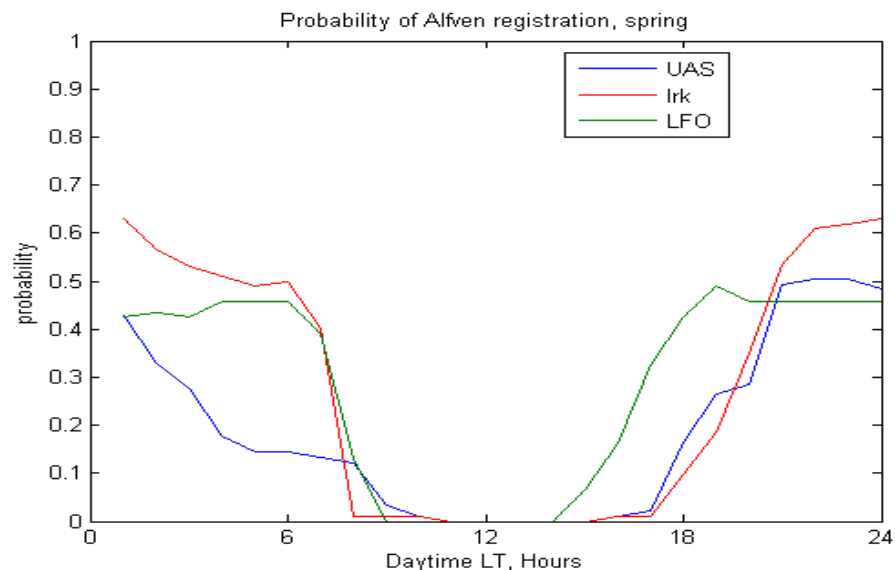
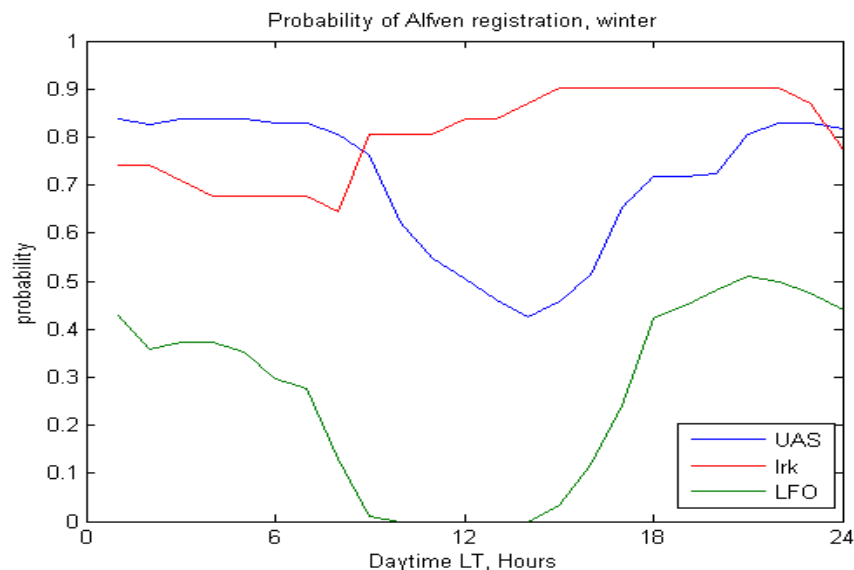
IAR Processing technique



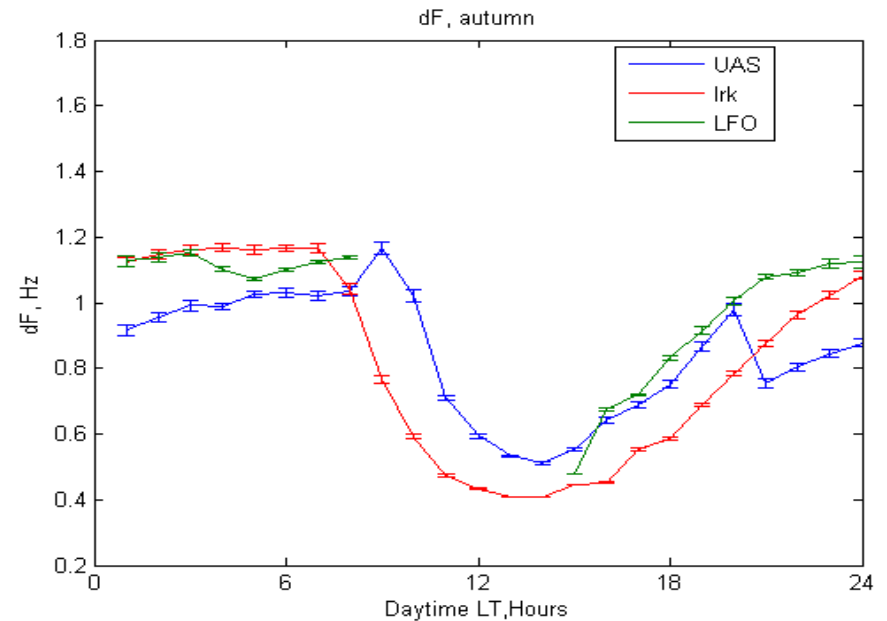
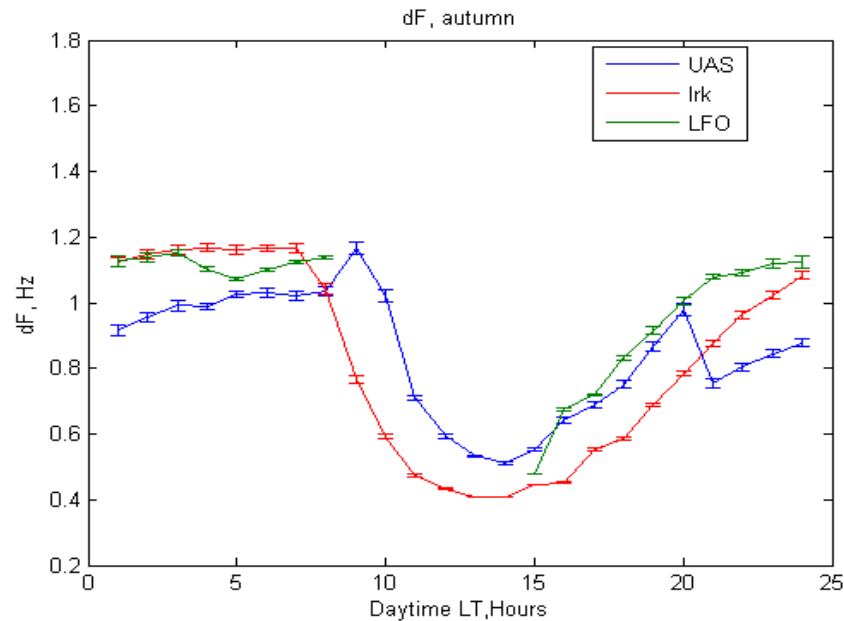
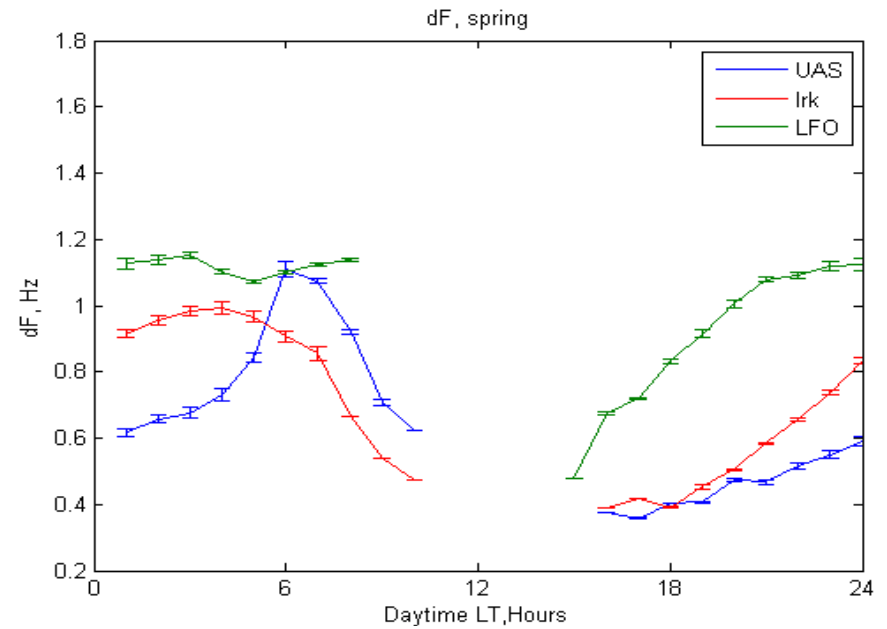
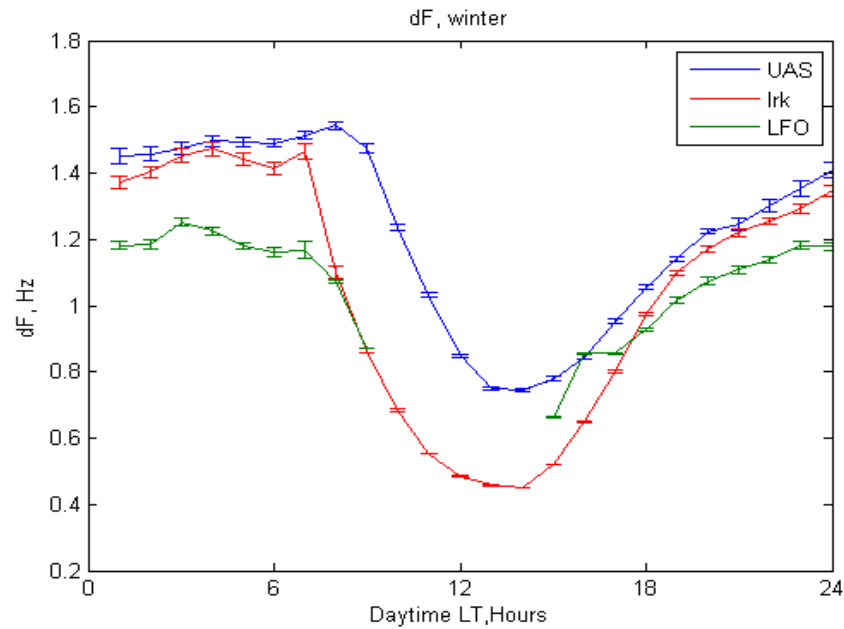
Recovered characteristics of IAR

- Probability of registration;
- Frequency spacing of RSS maximums dF

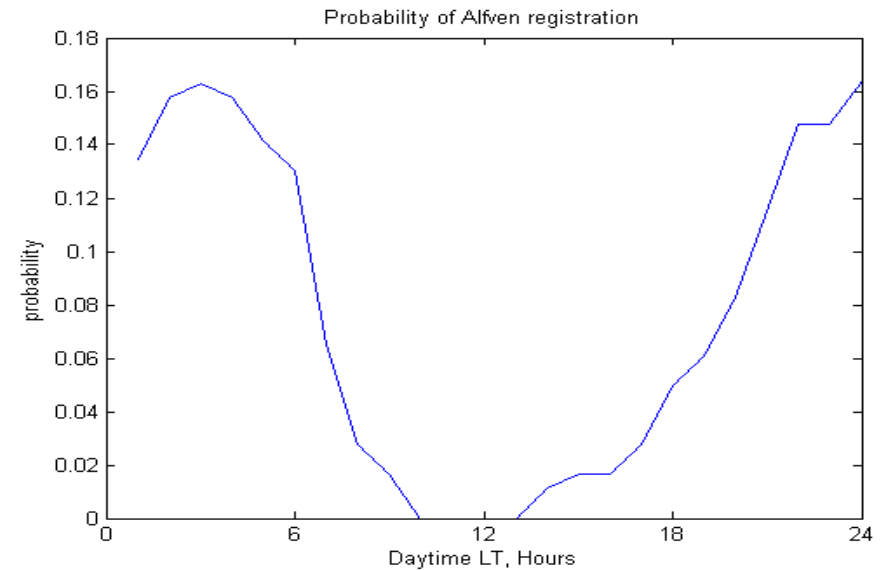
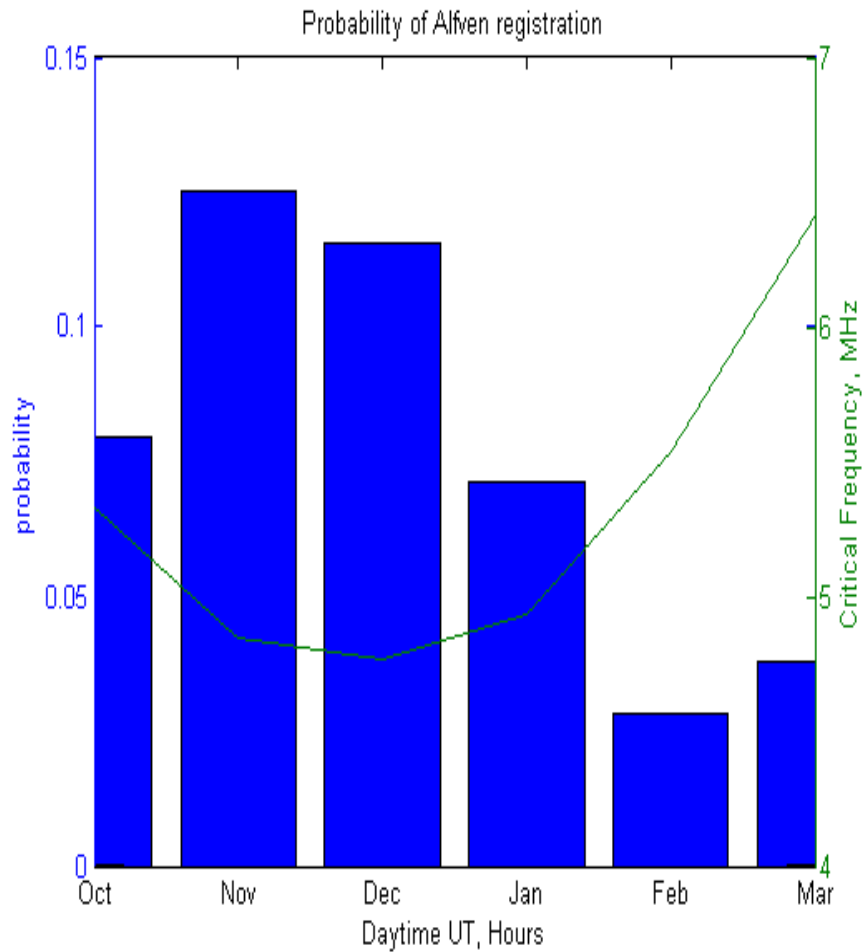
Seasonal-diurnal dependencies of probability of IAR registration



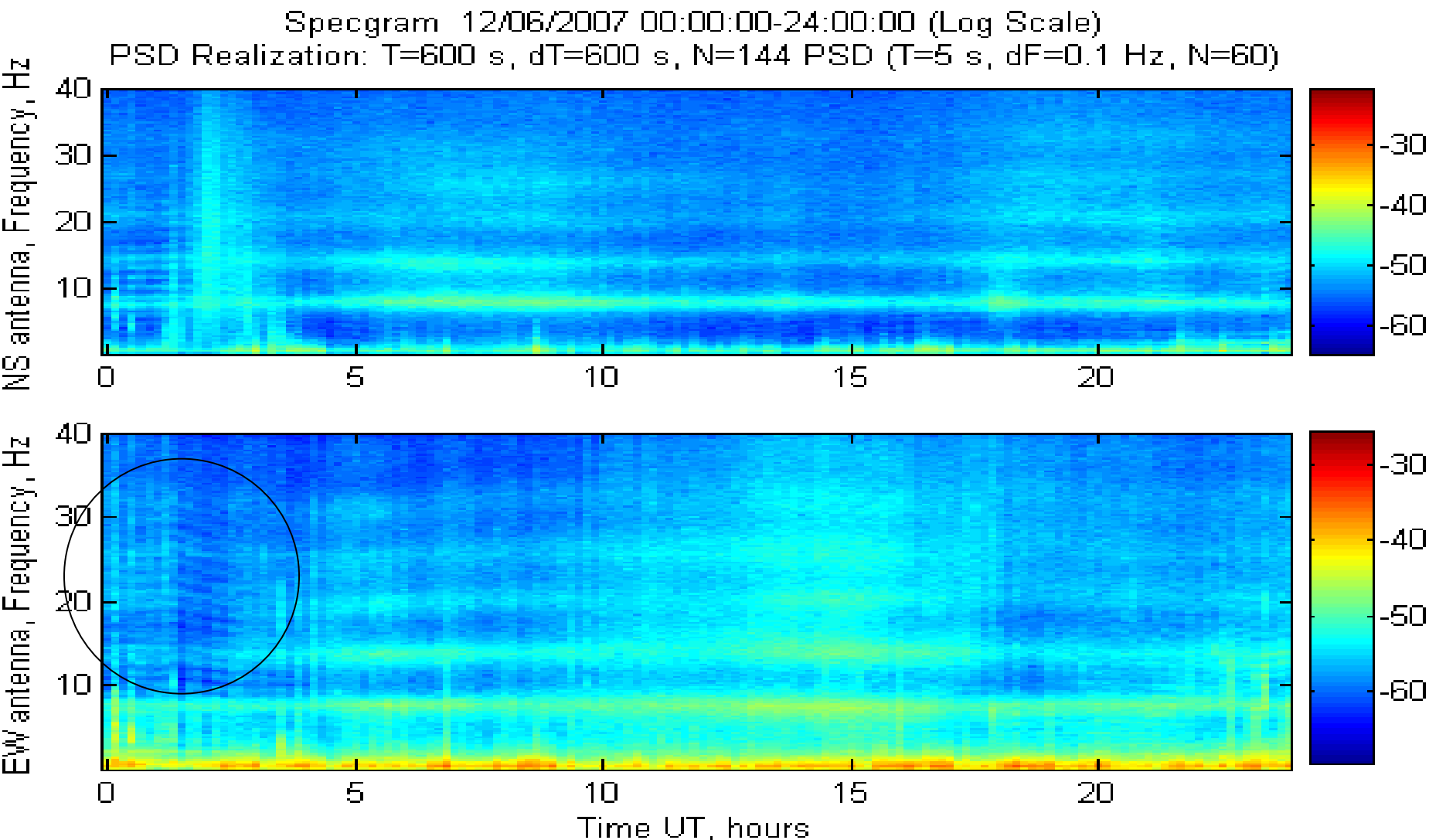
Seasonal-diurnal dependencies of dF



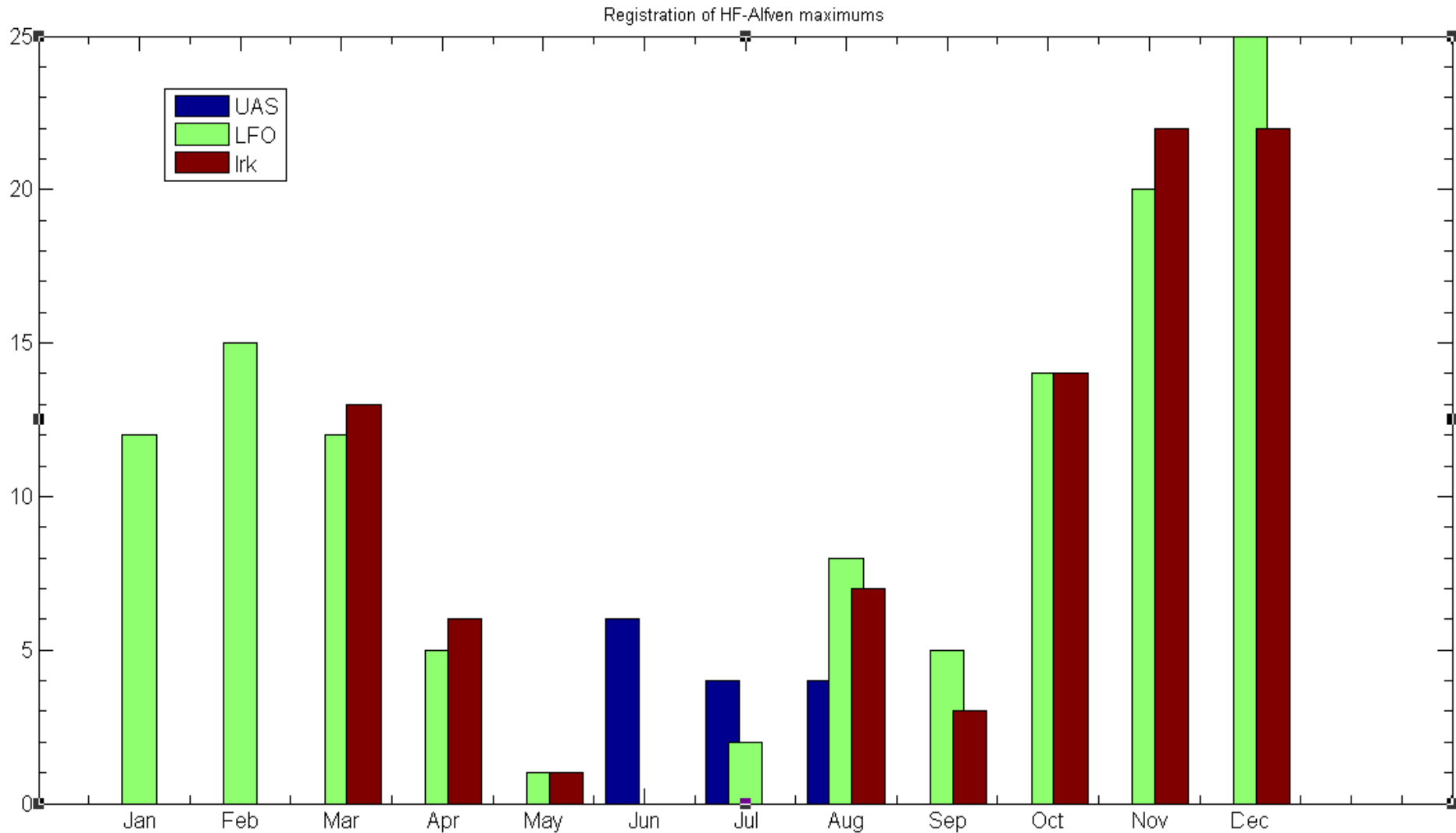
IAR data at SOUSY



Sample of registration of IAR maximums at the frequencies higher then 10Hz

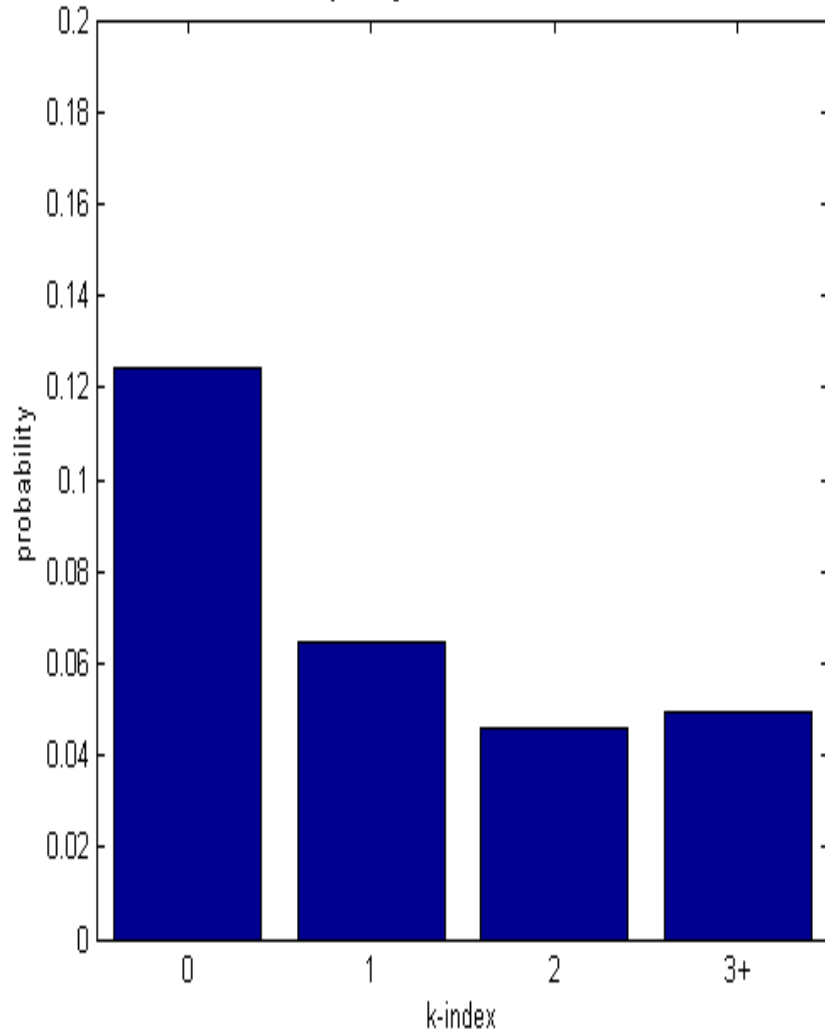


Statistic of registration of HF-IAR at different stations

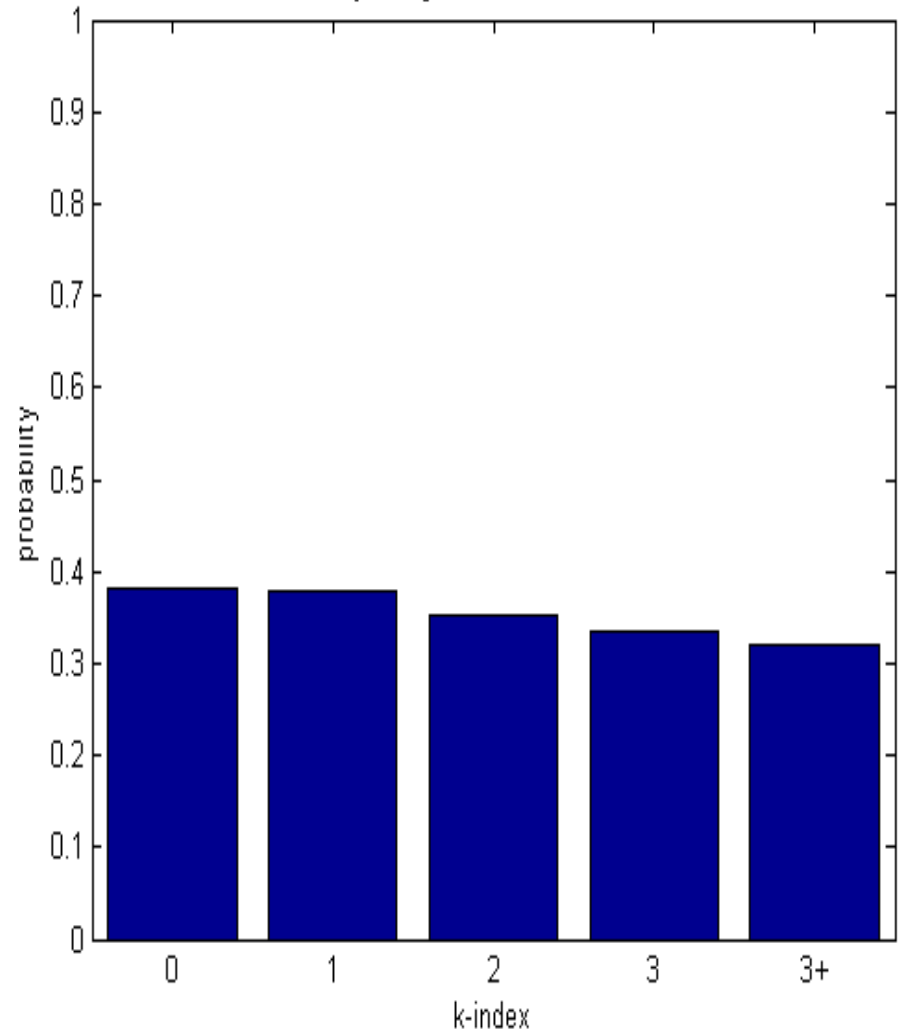


Relation between magnetic activity and probability of IAR registration

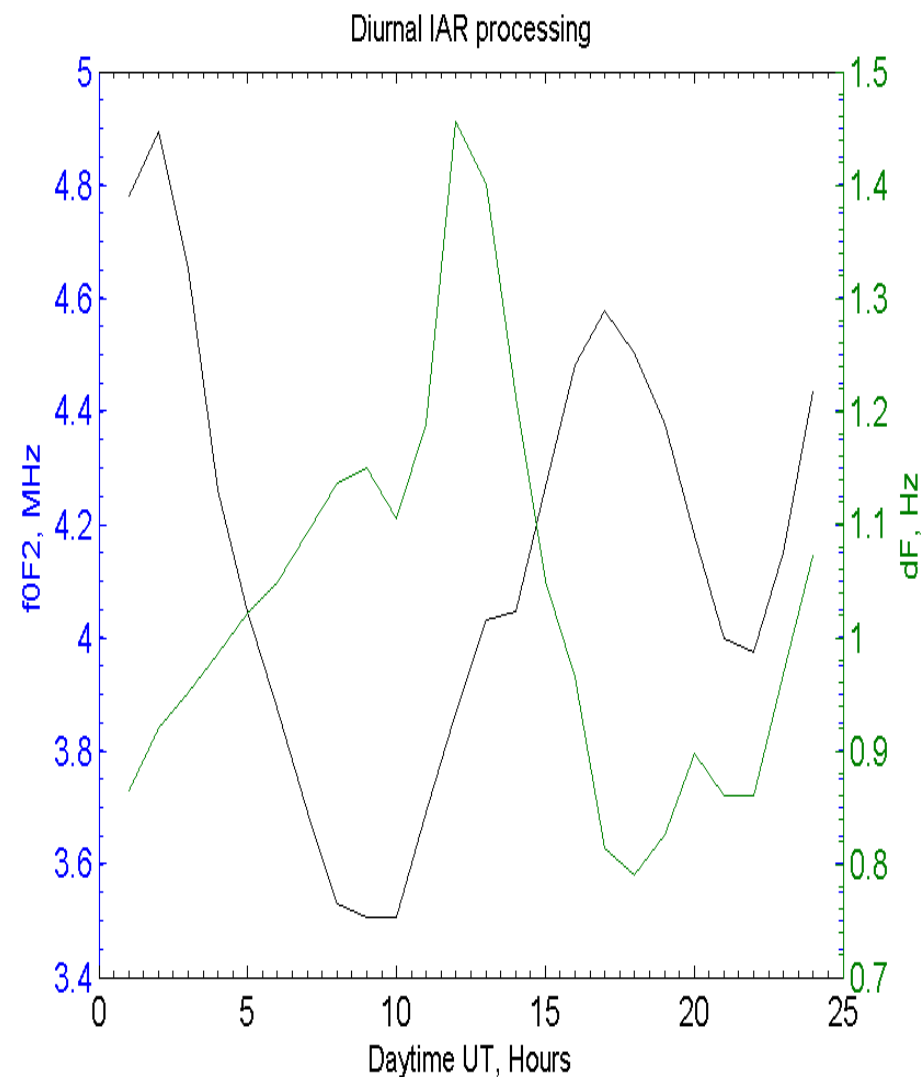
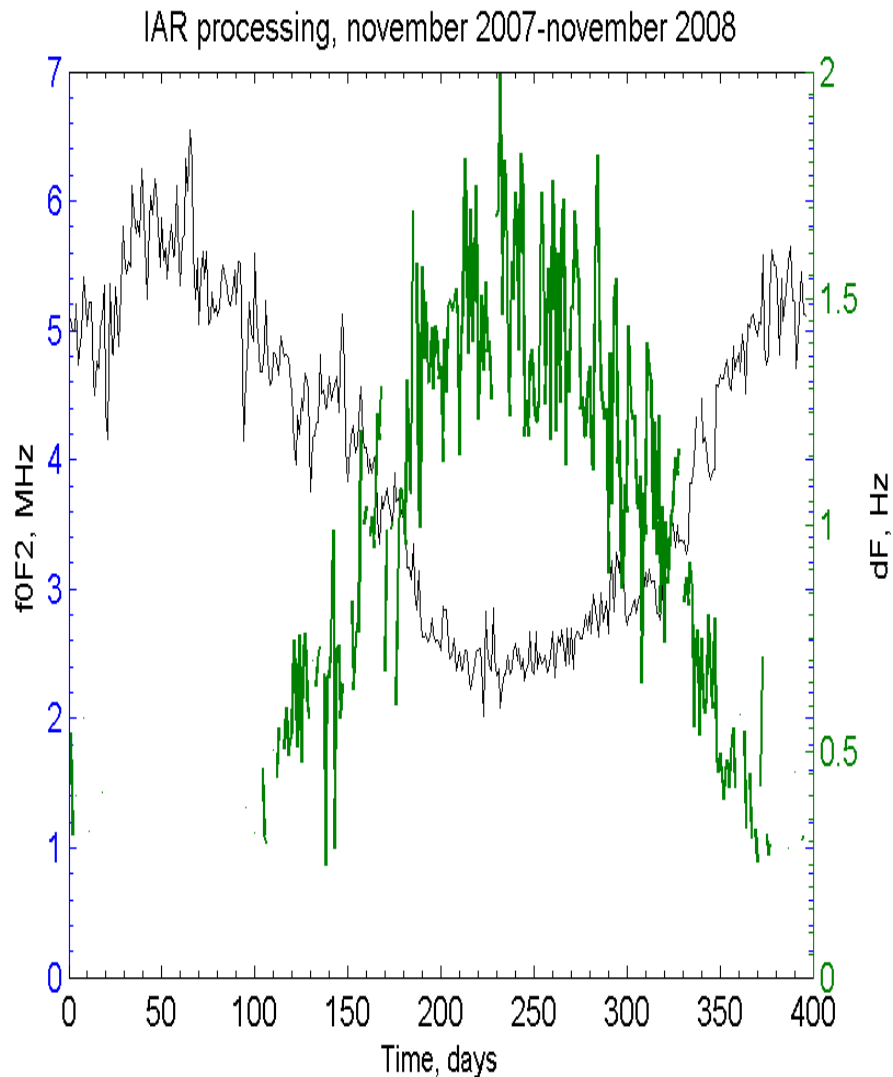
Probability of registration HF-IAR from k-index



Probability of registration of IAR from k-index

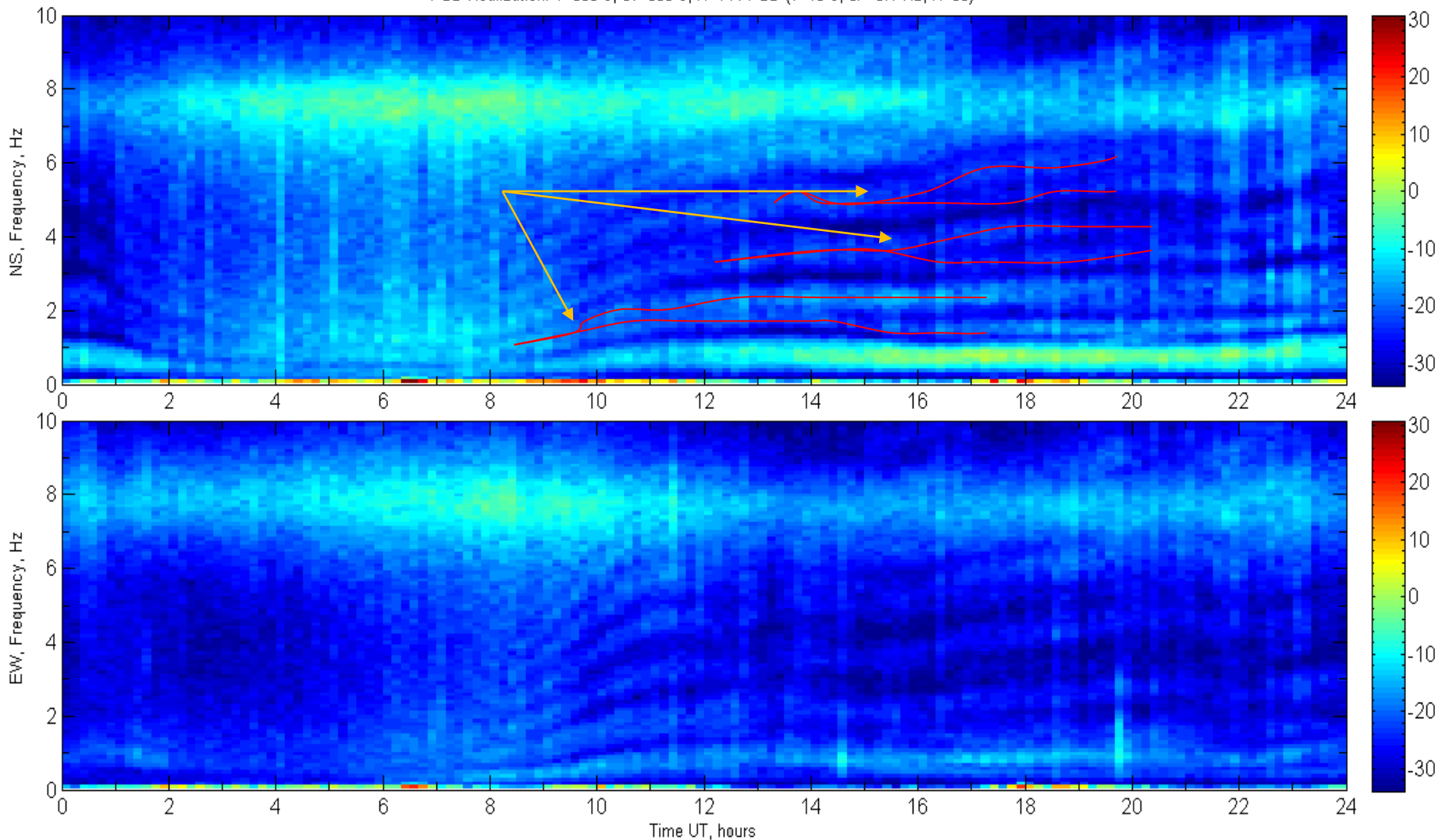


IAR frequency spacing and critical frequency

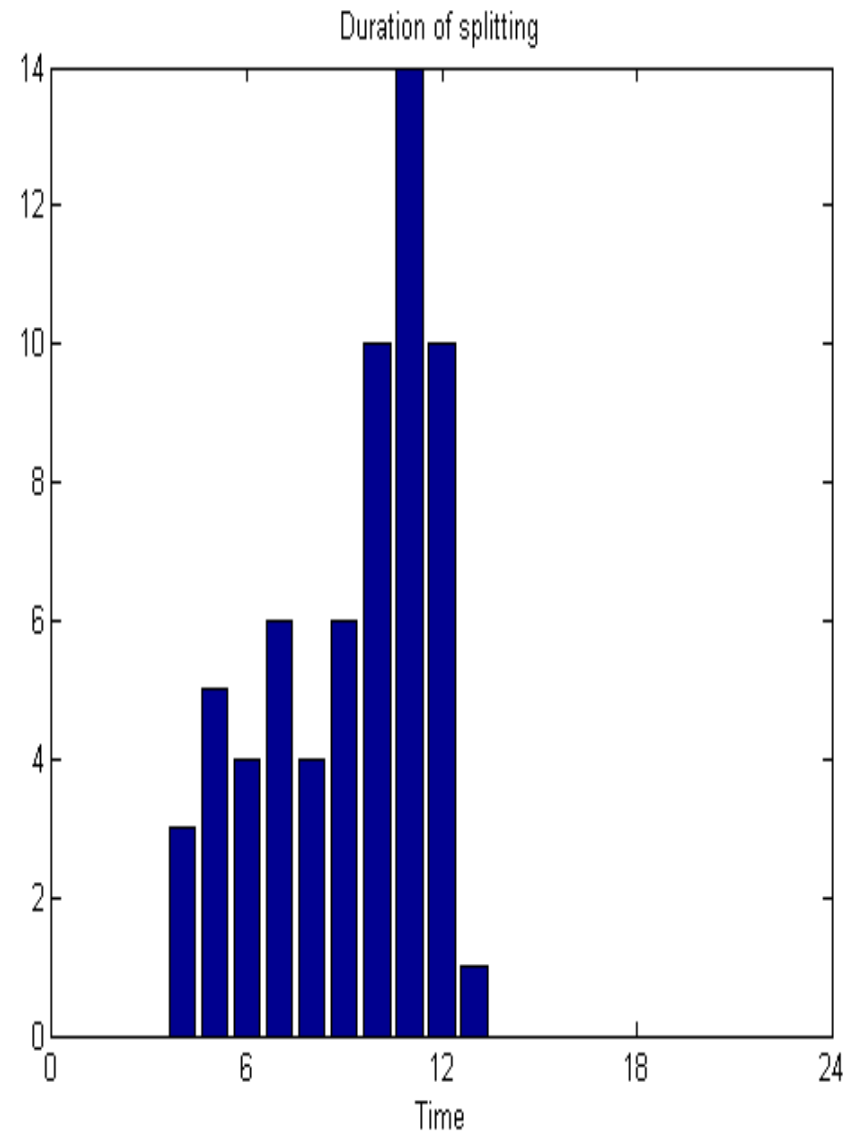
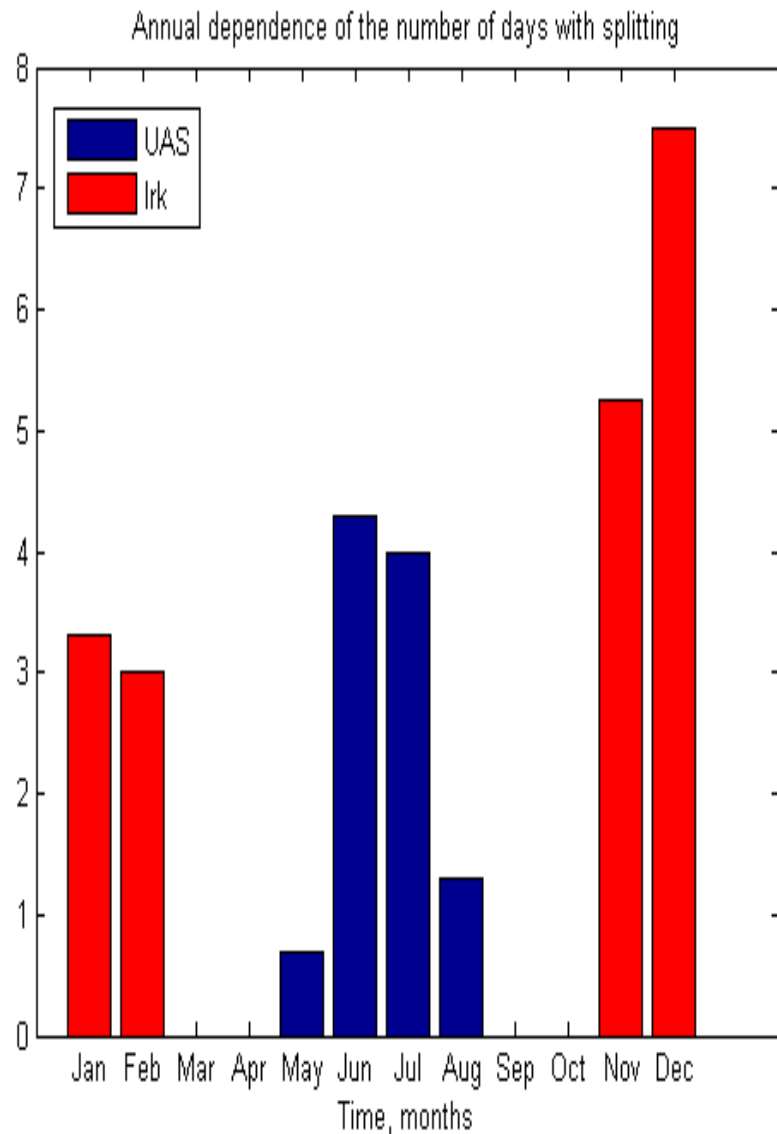


Splitting of the IAR maximums

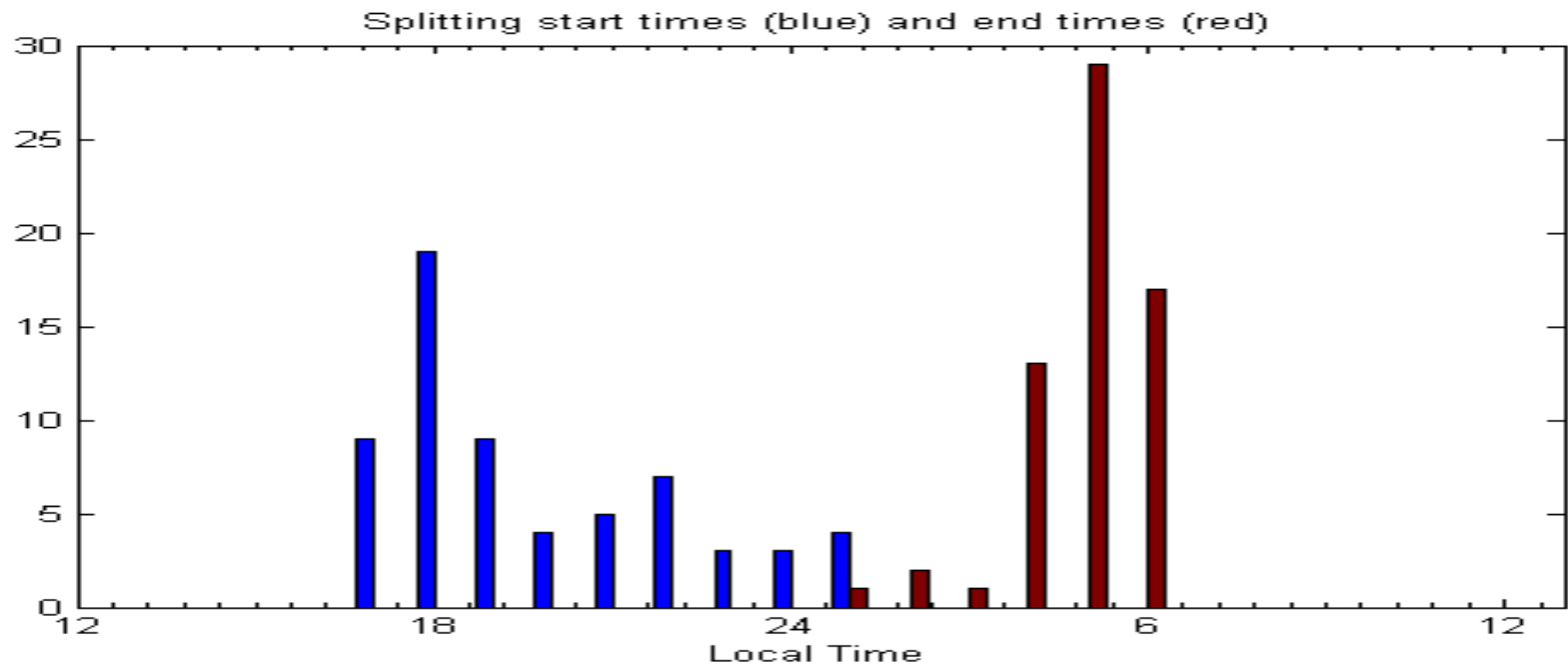
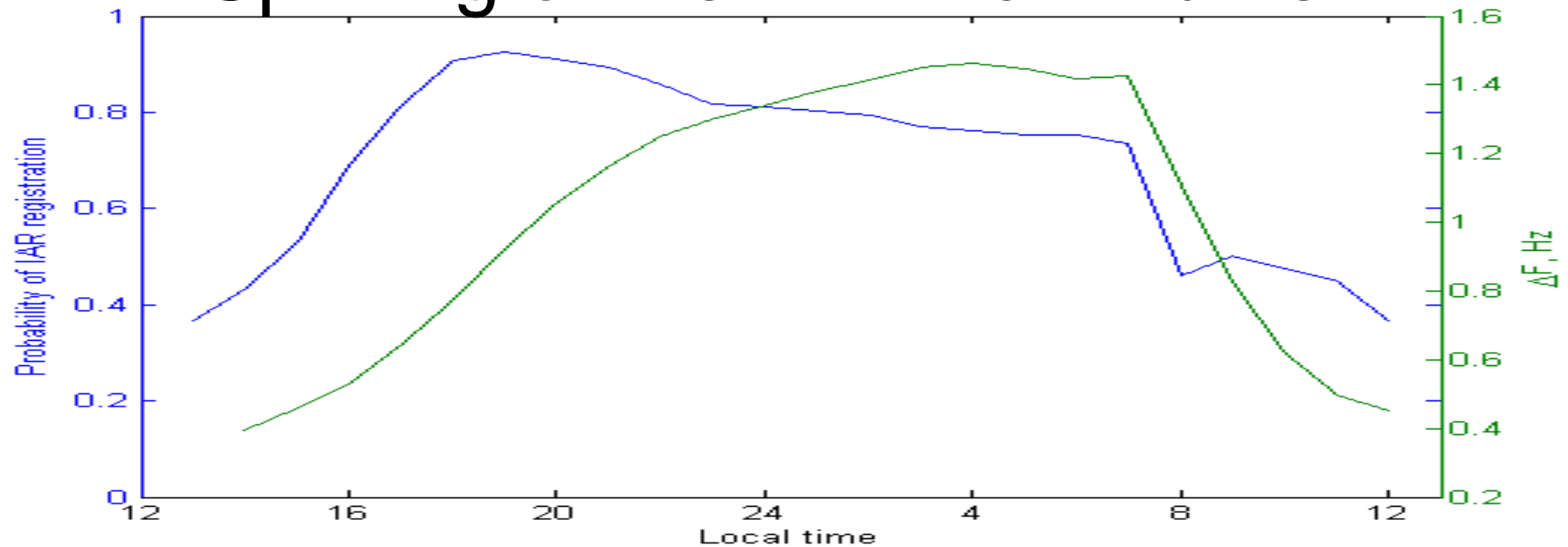
Specgram () 12/10/2010 00:00:00-24:00:00 UT (Log Scale),
PSD Realization: T=600 s, dT=600 s, N=144 PSD (T=10 s, dF=0.1 Hz, N=60)



Time dependences

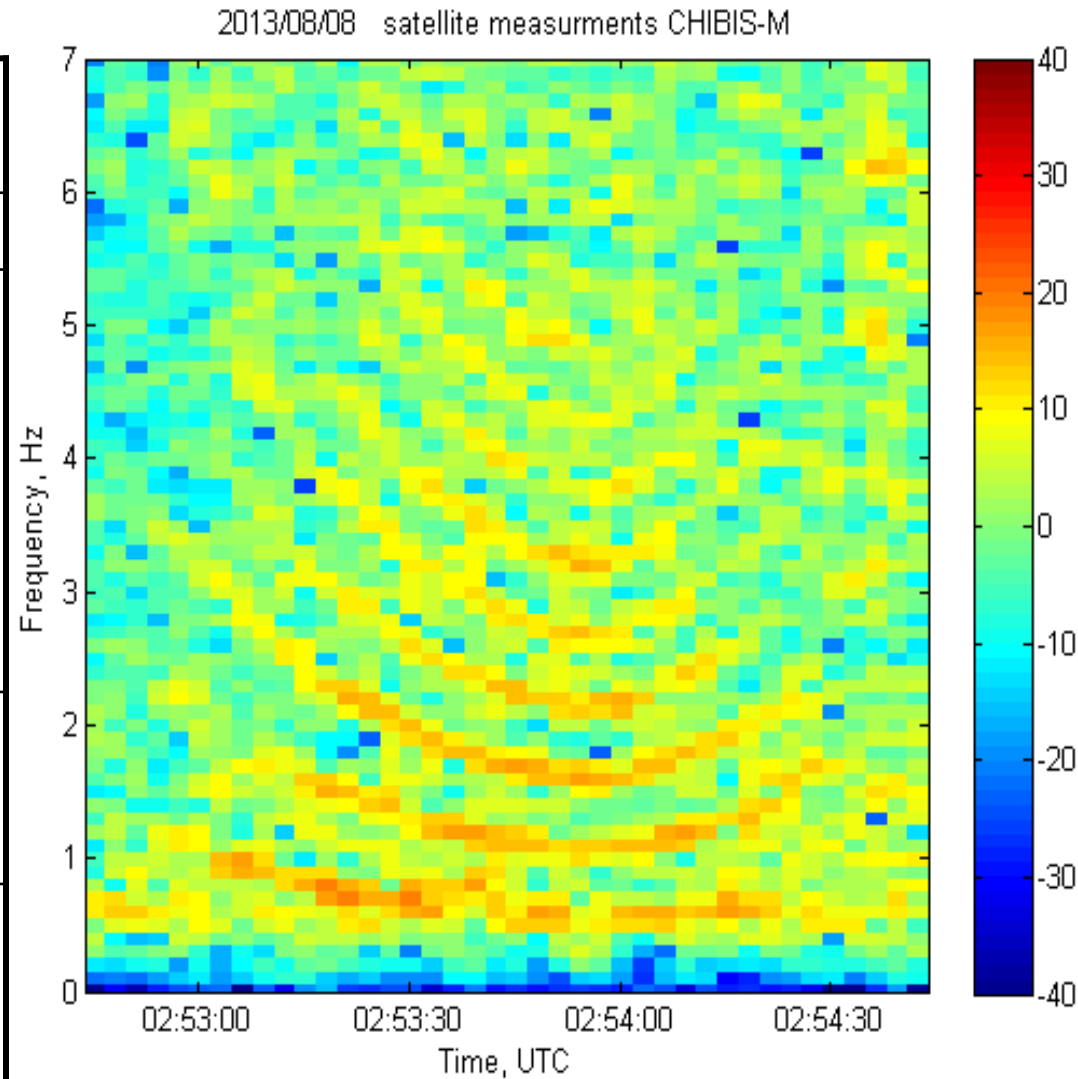


Splitting of the IAR maximums



Satellite Chibis-M

Frequency range for all channels	0.1 - 40 Hz
Dynamic range	120 dB
Noise level of measuring channels at 1 kHz	- Electric current density channel: $\leq 10\text{-}13 \text{ A/cm}^2 \text{ Hz}^{-1/2}$ - Electrical potential channel: $\leq 10\text{-}6 \text{ V Hz}^{-1/2}$
Current density transformation factor	77 V/mA/cm ²
Magnetic induction transformation factor	100 mV/nT
Power consumption	< 0.25 W

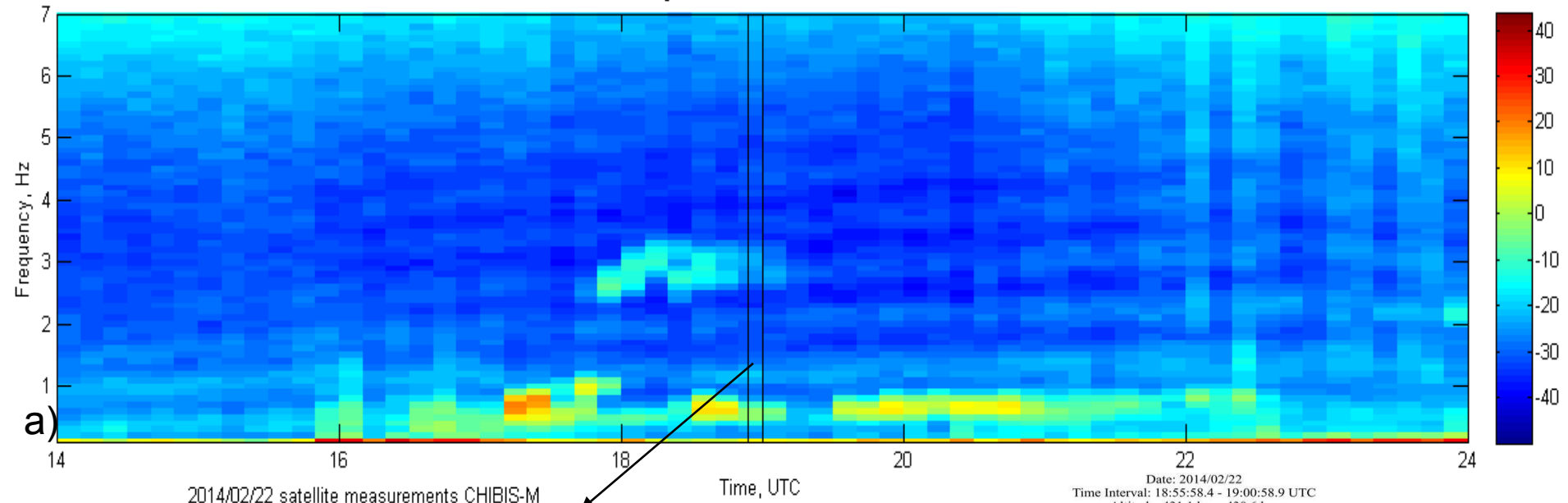


*D.Dudkin et al. Electric field signatures of the IAR and Shumann resonance in the upper ionosphere detected by Chibis -M microsatellite. JASTP 117 (2014)

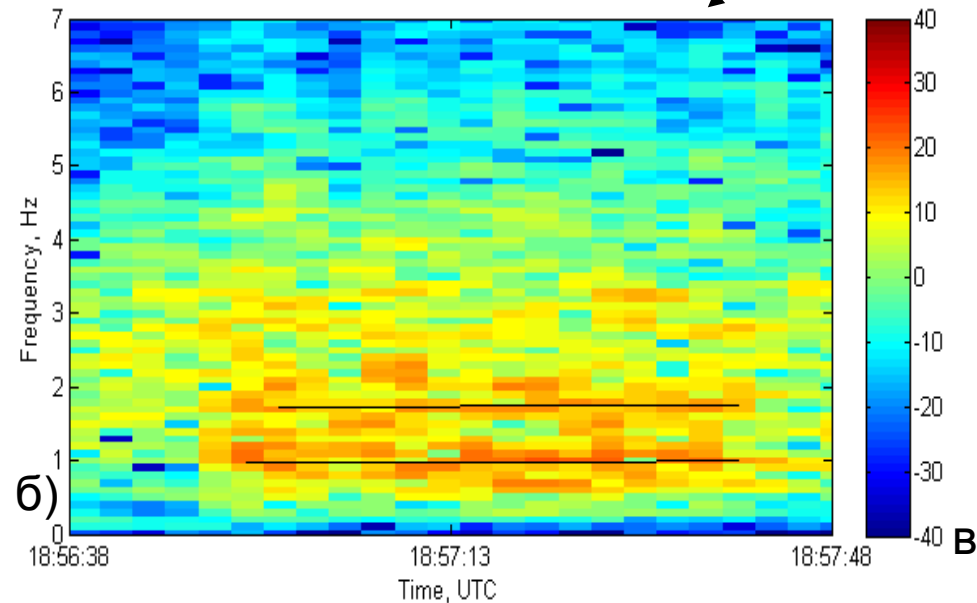
Ground based and satellite measurements

$\Delta F = 0.76$ Hz

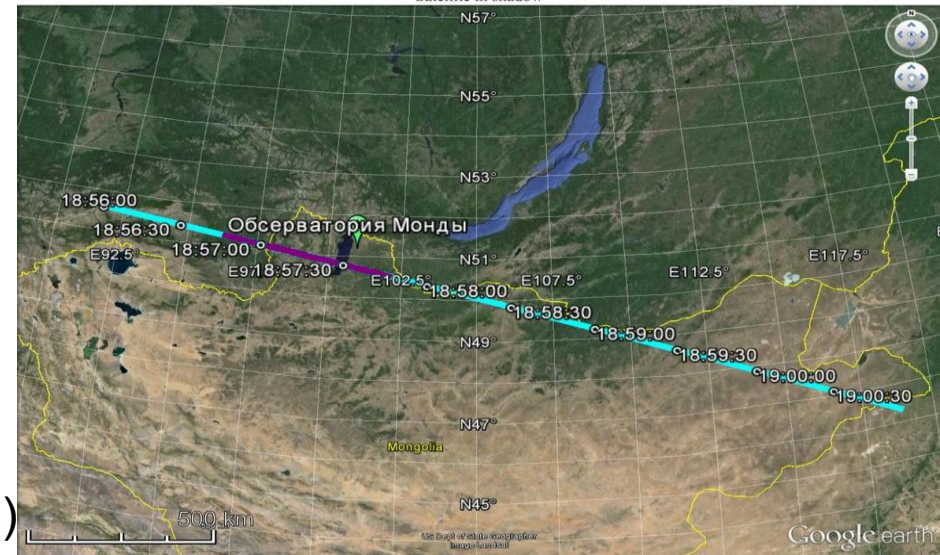
2014/02/22 ground-based measurements, Irkutsk



2014/02/22 satellite measurements CHIBIS-M



Date: 2014/02/22
Time Interval: 18:55:58.4 - 19:00:58.9 UTC
Altitude: 431.1 km - 430.6 km
Satellite in shadow



Conclusions:

- Diurnal and seasonal dependences of the IAR characteristics were recovered for the stations of observation spaced far away. We made comparative analysis of their morphology.
- SRS at frequencies up to 40 Hz were detected. We showed that in 80% of time such effect was observed not synchronous at different stations. Probability of such effect registration increases for quiet magnetic conditions.
- Effect of the splitting IAR maximums on two satellites was detected. The morphology of this phenomenon is analyzed.
- Experiments for the synchronous IAR observation in the space and on the Earth's ground were made. SRS was synchronously detected at the satellite and by ground base magnetometer or synchronously undetected. The frequency spacing is the same in the space and on the ground.

Thank You

Magnitude of the splitting

